

#### Jovian Planet Chemistry

Past the asteroid belt, there is a region of different planetary composition. Beyond about 4 AU from the Sun, water ice was able to condense and thus to become available as a raw material, in addition to the silicates and metals present in the inner Solar System.

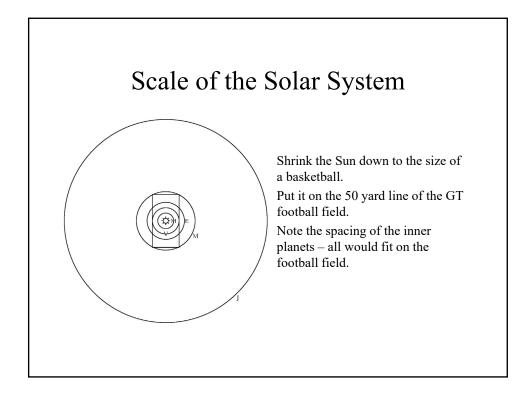
Since the atoms that constitute water are among the most abundant, a great deal of water ice is formed. [H is most abundant, O is 3<sup>rd</sup> most abundant.]

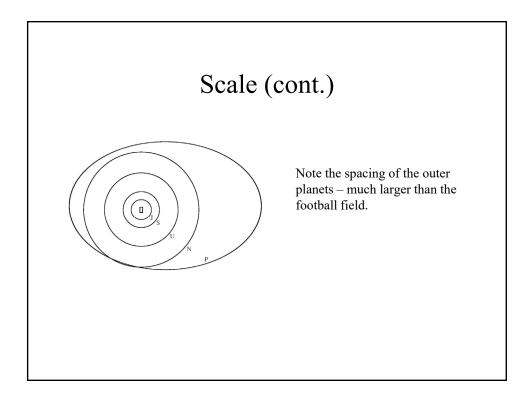
#### Jovian Planet Chemistry

The second major chemical distinction in the outer Solar System is the result of the **larger spacing** between planets and the accumulation of **more massive cores** of rock and ice.

The developing cores of Jupiter and Saturn grew large enough before the dissipation of the gaseous solar nebula to attract and hold the hydrogen and helium from large volumes of space.

Uranus and Neptune captured much less hydrogen and helium; this is why these two planets are both smaller than, and different in composition from, Jupiter and Saturn.





#### Jovian Planet Chemistry

With so much hydrogen available, the chemistry of the outer Solar System is **reducing**. Most of the oxygen present is chemically combined with hydrogen to make water, and it is therefore unavailable to form many oxidized compounds with other elements.

The compounds detected in the atmospheres of the giant planets are thus hydrogen-based gases, such as methane  $(CH_4)$  and ammonia  $(NH_3)$ .

## Exploration

**Pioneers 10 and 11** in 1973 & 1974. Primarily took a few pictures and measured magnetic fields.

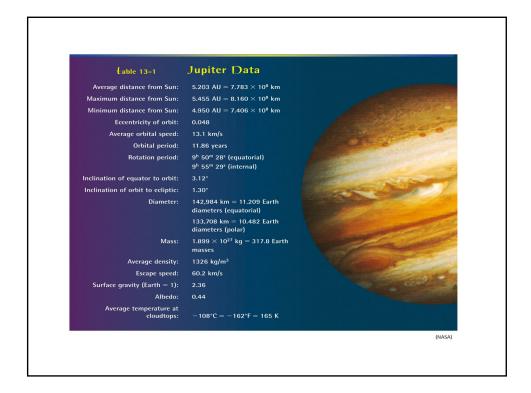
Voyagers 1 and 2 were in 1979. The Voyagers carried 11 instruments.

**Galileo** operated from 1995 to 2003. Its probe entered the Jovian Atmosphere.

New Horizons (Pluto) flyby in 2007.

Juno arrived in 2016.







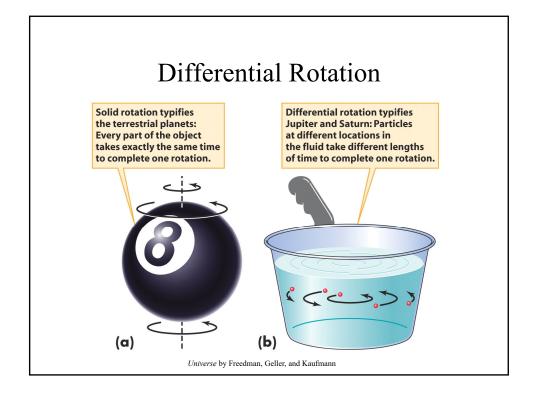


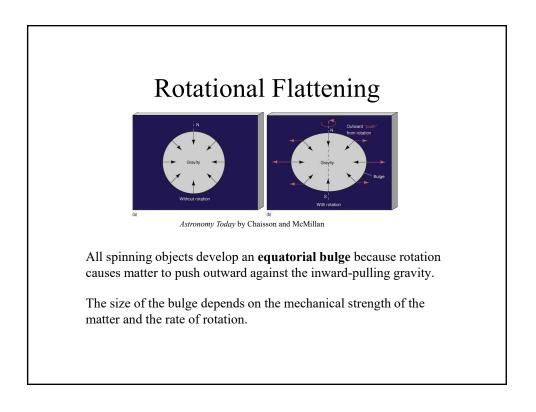
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Jupiter is a colorful and dynamic planet. Distinct details in its cloud patterns allow us to determine the rotation rate of the atmosphere.

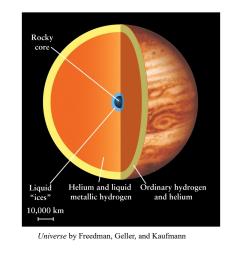
Much more fundamental is the rotation of the mantle and core, as indicated by periodic variations in the magnetic field. **This period of**  $9^{h}56^{m}$  gives Jupiter the shortest day of any planet.

However, Jupiter experiences differential rotation.





# Composition and Structure

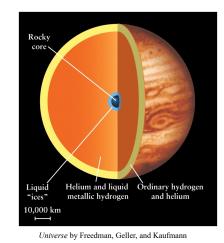


The internal structure is different than that of the terrestrial planets.

At depths of only a few thousand km, pressures become so high that hydrogen changes from **gaseous** to a **liquid** state.

Still deeper, this liquid hydrogen can act like a **metal**. The greater part of the interior is **liquid metallic hydrogen**.

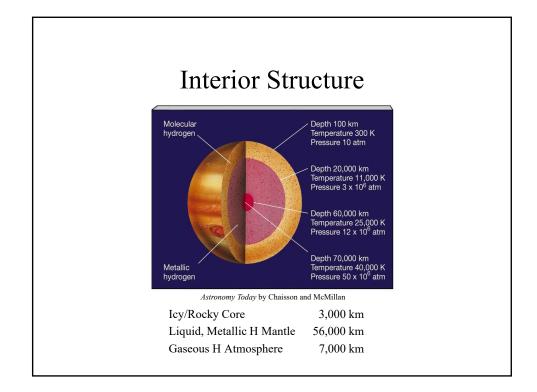
#### Composition and Structure

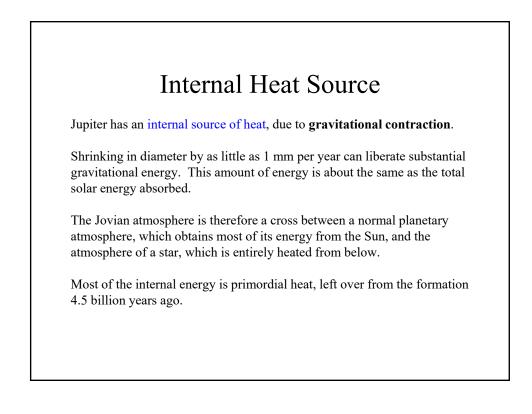


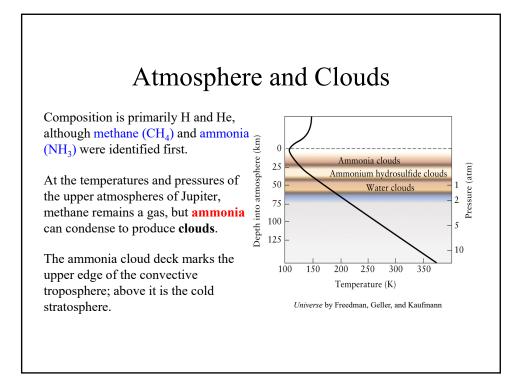
The core is composed of heavier materials, presumably the original rock-and-ice small bodies.

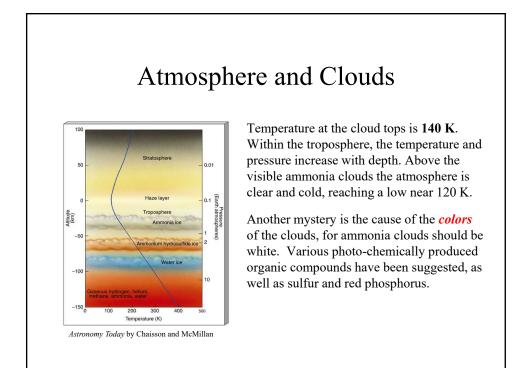
Models predict a central pressure of over 100 million bars and a central density of about  $31 \text{ g/cm}^3$ .

[Much greater than the Earth's, but not nearly enough to be a star.]

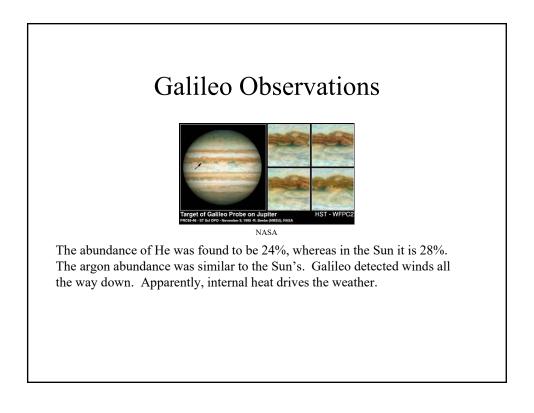








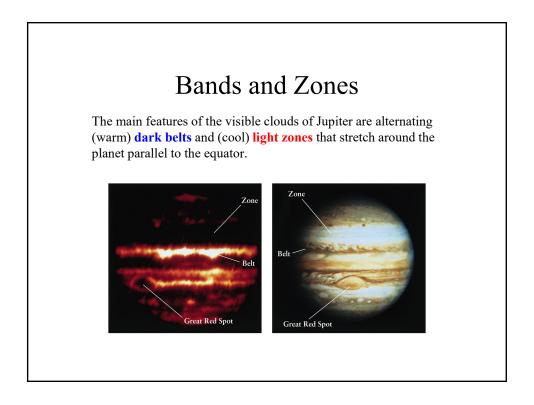
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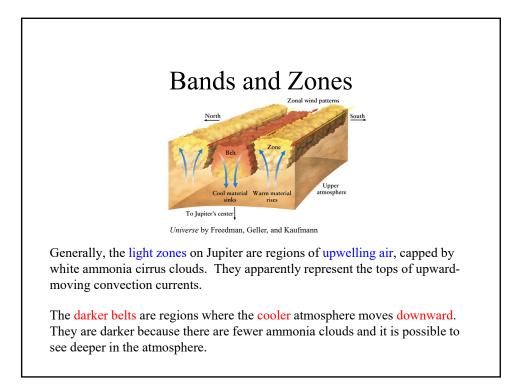


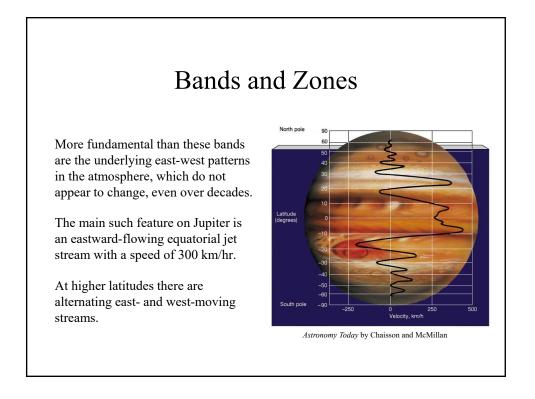
### Winds and Weather

The atmospheric dynamics observed on the Jovian planets differ from those of the terrestrial planets. The three primary reasons are:

- 1. These planets have much deeper atmospheres, with no solid lower boundary,
- 2. They spin faster than the terrestrial planets, suppressing northsouth circulation patterns and accentuating east-west airflow, and
- 3. On all except Uranus, internal heat sources contribute about as much energy as sunlight, forcing the atmospheres into deep convection to carry the internal heat outward.







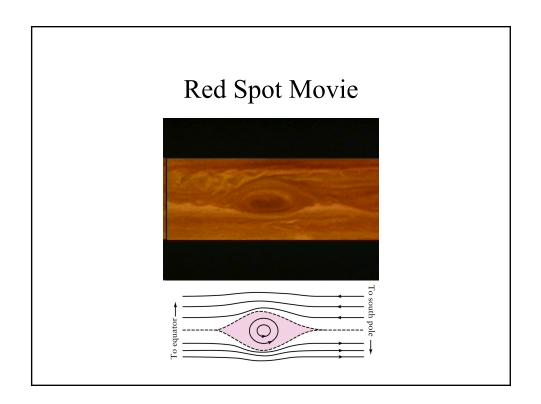
#### Storms

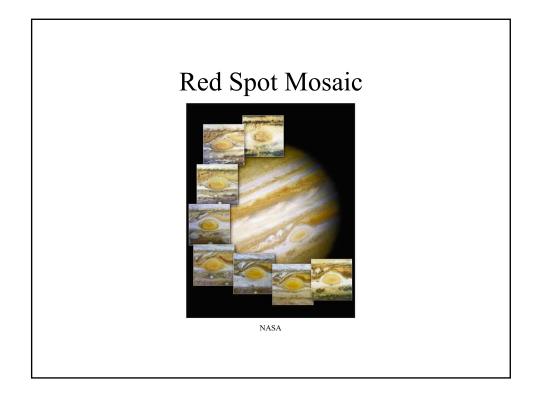


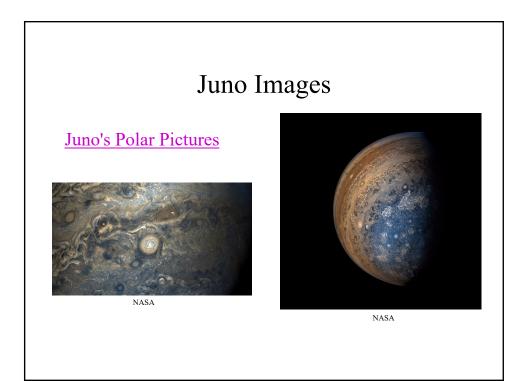
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The most famous storm on Jupiter is the **Great Red Spot**, a reddish oval that is big enough to hold two Earths side by side. It has existed for over 300 years.

There are also three smaller, white ovals, which formed in the 1940s. The cause is unknown. They last long times because there are no solid surfaces to slow down atmospheric disturbances. Furthermore, their large size lends to stability.





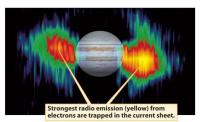


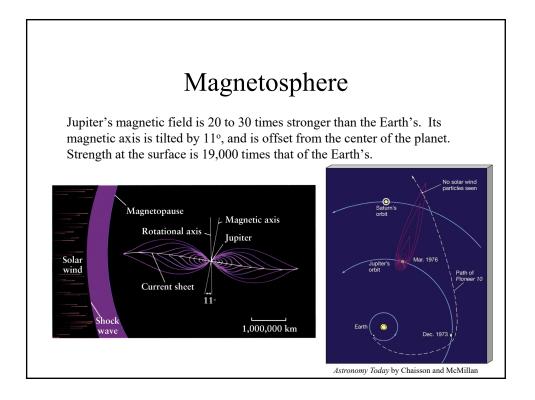
## Magnetosphere

In the 1950s, radio energy was observed from Jupiter that was more intense at longer than at shorter wavelengths – just the reverse of what is expected from thermal radiation. This is typical of the radiation emitted by electrons accelerated by a magnetic field, called **synchrotron radiation**.

Later observations showed that the radio energy originates from a region surrounding the planet whose diameter is several times that of Jupiter itself.

The evidence suggests that there are a vast number of charged atomic particles circulating around Jupiter, spiraling through the lines of force of a magnetic field associated with the planet.



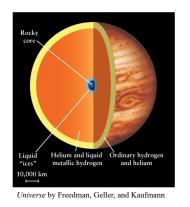


## Magnetosphere

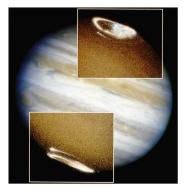
Presumably the magnetic fields of the outer planets are generated in much the same way as the field of the Earth.

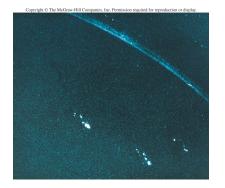
All of these planets **spin rapidly**, so there is a ready source of energy to power their internal magnetic generators.

Jupiter has a large interior region of metallic liquid hydrogen that acts like the liquid iron core of the Earth.



### Aurorae and Lightning





Universe by Freedman, Geller, and Kaufmann

The oval-shaped aurorae, extending hundreds of km above Jupiter's limb, are created by charged particles escaping from the magnetosphere and colliding with the atmosphere, which causes the gas to glow.

## Rings

A ring is a collection of vast numbers of particles -a ring is not a solid "washer". Each particle obeys Kepler's Laws as it follows its own orbit around the planet. Thus the inner particles orbit faster than those farther out, and the ring as whole does not rotate as a solid body.

Jupiter's ring was discovered by Voyager 1 in a single image that was targeted specifically to search for a faint ring system. Subsequently, Voyager 2 was reprogrammed to take a more complete set of images.

### Rings of Jupiter

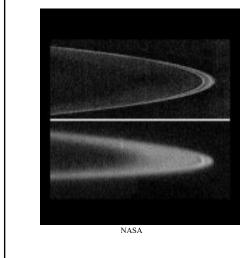


NASA

The ring is now known to be composed of three major components. The "Main" ring is about 7,000 km wide. The main ring encompasses the orbits of two small moons, Adrastea and Metis, which may act as the source for the dust that makes up most of the ring.

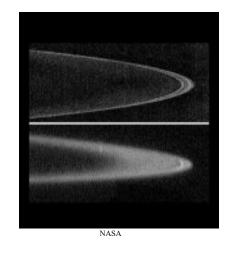
At its inner edge the main ring merges gradually into the "Halo." The halo is a broad, faint torus of material about 20,000 km thick and extending halfway from the main ring down to the planet's cloudtops.

# Newest Ring Images



The New Horizons spacecraft took the best images of Jupiter's charcoal-black rings as it approached and then looked back at Jupiter. The image is sharply focused, though it appears fuzzy due to the cloud of dust-sized particles enveloping the rings.

### Newest Ring Images



Just outside the main ring is the broad and exceedingly faint "Gossamer" ring, which extends out beyond the orbit of the moon Amalthea.

It is probably composed of dust particles less than 10 microns in diameter – about the size of cigarette smoke particles.

The origin of the ring is probably from micrometeorite bombardment of the tiny moons orbiting within the ring.